

The precise determination of the  $^{13}\text{C}$  isotopic ratios for carbon chain molecule  $\text{HC}_3\text{N}$  in the low-mass star forming region L1527 by radio observations

\*Mitsunori Araki<sup>1</sup>, SHURO TAKANO<sup>2</sup>, NAMI SAKAI<sup>3</sup>, SATOSHI YAMAMOTO<sup>4</sup>, TAKAHIRO OYAMA<sup>1</sup>, NOBUHIKO KUZE<sup>5</sup>, KOICHI TSUKIYAMA<sup>1</sup>

1.Infrared Free Electron Laser Research Center, Tokyo University of Science, 2.Nihon University, 3.The Institute of Physical and Chemical Research (RIKEN), 4.The University of Tokyo, 5.Sophia University

Linear carbon chains have been found in dark clouds. Recently a lot of carbon chains have also been detected in the class 0 object L1527, which is a low-mass star forming region. We observed the normal and isotopic species of the fundamental carbon-chain molecule  $\text{HC}_3\text{N}$  in L1527 with the high signal-to-noise ratios using Green Bank 100 m telescope and Nobeyama 45 m telescope to explore the production scheme of  $\text{HC}_3\text{N}$  in the region having a warm carbon chain chemistry (WCCC). The spectral lines of the rotational transitions in the 44–109 GHz region were used to study the abundances of the normal and isotopic species. The ratios were precisely determined to be  $1.00 : 1.01 \pm 0.03 : 1.35 \pm 0.04 : 86.4 \pm 2.2$  for  $[\text{H}^{13}\text{CCCN}] : [\text{HC}^{13}\text{CCN}] : [\text{HCC}^{13}\text{CN}] : [\text{HCCCN}]$ , where the errors are in one standard deviations. The ratios of  $[\text{HCC}^{13}\text{CN}]/[\text{HCCC}^{15}\text{N}]$  and  $[\text{HCCCN}]/[\text{HCCC}^{15}\text{N}]$  were obtained to be  $5.26 \pm 0.24$  and  $338 \pm 15$ , respectively. It was found that the abundance of  $\text{H}^{13}\text{CCCN}$  is almost equal to that of  $\text{HC}^{13}\text{CCN}$ , and it was concluded that  $\text{HC}_3\text{N}$  in L1527 is mainly formed by the reaction schemes *via*  $\text{C}_2\text{H}_2$  and  $\text{C}_2\text{H}_2^+$ , *i.e.* dicarbide chemistry. The obtained ratios are quite similar to those reported in the cyanopolyne peak in the starless dark cloud Taurus Molecular Cloud-1 [2]. Thus this similarity would suggest a universality of dicarbide chemistry producing  $\text{HC}_3\text{N}$  irrespective of evolutionary phases from a starless dark cloud to a class 0 object for a low-mass star forming region.

[1] Sakai, N., Sakai, T., Hirota, T., & Yamamoto, S. 2008, ApJ, 672, 371

[2] Takano, S., Masuda, A., Hirahara, Y., et al. 1998, A&A, 329, 1156

Keywords: radio, carbon chain molecule, isotopic ratio