Compositional fluctuation of silicate minerals in formation processes of artificial cosmic spherules

*Hiroshi Isobe¹, Minami Ohno²

¹Department of Earth and Environmental Sciences, Faculty of Advanced Science and Technology, Kumamoto University, ²Department of Earth and Environmental Sciences, Faculty of Science, Kumamoto University

Micrometeorites (MMs) have the largest flux of planetary materials falling to the earth. Cosmic spherules (CSs) are spherical MMs melted by heating at the atmospheric entry. CSs experienced extraordinary quick quench processes following various degree of melting. In this study, quick heating and cooling experiments of powdered meteorite were carried out with a fine particles free falling apparatus with controlled gas flow (Isobe and Gondo, 2013). Oxygen fugacity was controlled to FMQ +1.5 log unit to represent approximately 90 km altitude in the terrestrial atmosphere. Particles can be heated to the maximum temperatures of 1400°C or 1500°C within two seconds, are kept approximately one second and quenched within a second.

Run products are quite resemble to natural CSs. Olivine crystals in meteorite with various Mg/Fe compositions in Fo60 to 90 were melted depending on the phase relation of olivine solid solution. Compositions and texture of recrystallized olivine show crystallization process in quick quench. Forsterite component if recrystallized olivine is concentrated to Fo 70 to 85. Pyroxenes in meteorites show also various compositions including orthopyroxene, pigeonite and augite. However, pyroxene phenocrysts in reproduced CSs are only relict enstatite. Pyroxene quench crystals are scarcely occur in the run products. Plagioclase phenocrysts with Na-rich composition (Ab 78 to 98) in meteorites were also melted by heating. Plagioclase phenocrysts in run products show various compositions in Ab 10 to 60.

Melting of silicate minerals follow phase relations of each mineral species even in extraordinary quick heating and short durations of melt. Compositions of olivine growing in the quick recrystallization process show strong disequilibrium. Recrystallized plagioclase composition is enriched in Ca due to Ca supplied by melting of clinopyroxene. Textures of CSs are depending on melting degree of the source materials. Mineralogical texture is also depending on thermal history and oxidation by atmosphere. Compositions of recrystallized phenocrysts in CSs should be resulted in quick quenching processes.

Keywords: micrometeorite, cosmic spherule, quench crystal, atmospheric heating