

## ペルム紀／三疊紀境界における地球外<sup>3</sup>Heフラックスの増加イベント Enhanced flux of extraterrestrial <sup>3</sup>He across the Permian–Triassic boundary

\*尾上 哲治<sup>1</sup>、高畑 直人<sup>2</sup>、佐藤 峰南<sup>4</sup>、石川 晃<sup>3</sup>、曾田 勝仁<sup>1</sup>、佐野 有司<sup>2</sup>、磯崎 行雄<sup>2</sup>

\*Tetsuji Onoue<sup>1</sup>, Naoto Takahata<sup>2</sup>, Honami Sato<sup>4</sup>, Akira Ishikawa<sup>3</sup>, Katsuhito Soda<sup>1</sup>, Yuji Sano<sup>2</sup>, Yukio Isozaki<sup>2</sup>

1. 熊本大学、2. 東京大学、3. 東京工業大学、4. 千葉工業大学

1. Kumamoto Univ., 2. Univ. Tokyo, 3. Tokyo Inst. Tech., 4. Chiba Inst. Tech.

The Solar System contains abundant submillimetre interplanetary dust particles (IDPs) that are enriched in <sup>3</sup>He. <sup>3</sup>He concentrations in ancient deep-sea sediments have been used to constrain the flux of IDPs onto the Earth for at least the last 100 Myr. However, the use of <sup>3</sup>He in detecting IDP flux is often compromised by the diffusional loss of <sup>3</sup>He in sedimentary rocks, with the exception of some Ordovician samples that record a period of unusually high extraterrestrial <sup>3</sup>He (<sup>3</sup>He<sub>ET</sub>) flux. In this study, we report for the first time the preservation of extra-terrestrial <sup>3</sup>He in deep-sea bedded chert from a continuous Permian-Triassic boundary (PTB) section at Waidani in Japan, which was deposited in the Panthalassa superocean.

High <sup>3</sup>He/<sup>4</sup>He ratios (up to 150 Ra; 1Ra = the atmospheric ratio) were detected from acid-insoluble residues from the uppermost Permian deposits, which suggest the existence of extraterrestrial He hosted mainly in IDPs. The estimated extraterrestrial fraction of <sup>3</sup>He across the PTB reveals that <sup>3</sup>He<sub>ET</sub> concentrations are higher in the topmost ~1.5 m of the studied Permian deposits, which is 4–5 times greater than that measured in the overlying Triassic unit. Based on the <sup>3</sup>He<sub>ET</sub> concentration and sedimentary mass accumulation rate of the PTB section at Waidani, we calculated <sup>3</sup>He<sub>ET</sub> flux across the PTB. The data document the presence of an up to 4-fold increase in IDP flux for the last 500 kyr interval of the Permian. This unusual signal suggests a significant increase in the influx of interplanetary dust particles, likely related to an asteroid shower in the inner Solar System. High-resolution stratigraphy indicates that peak IDP flux occurred during the final 500 kyr of the Permian, concurrent with a pre-extinction decline in radiolarian diversity.

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