

Stochastic Shock Drift Acceleration in the Shock Transition Region with Finite Pitch-Angle Anisotropy

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The acceleration of non-thermal electrons is one of the important subjects in space physics. The shock accelerated non-thermal electrons have been observed at Earth's bow shocks (Gostling et al. 1989). These electrons are supra-thermal (with energies from 1 keV to 100 keV), and their distribution is power-law. However, acceleration models proposed in the past could not reproduce such observed spectra. Recent in-situ satellite observations of Earth's bow shock found wave-particle interactions between whistler waves and electrons at the shock transition region (Oka et al. 2017). These results indicate that whistler waves play an important role for the acceleration of supra-thermal electrons.

We propose the stochastic shock drift acceleration (SSDA) model as a new acceleration mechanism for non-thermal electrons. It takes into account the effect of stochastic pitch-angle scatterings by whistler waves during the course of the Shock Drift Accelerations (SDA), which is a classical acceleration model for supra-thermal electrons at the shock transition region (Wu 1984, Leroy and Mangeney 1984). By introducing stochasticity with pitch-angle scatterings, the acceleration efficiency may be improved. We theoretically analyzed this acceleration model. We showed that the electron energy spectrum becomes a power-law consistent with observations in the limit of strong scatterings. We also found the maximum energy attainable through the proposed model, and it scales linearly with the pitch-angle scattering coefficient which represents the strength of scatterings. In our previous study, we have assumed the electron pitch-angle distribution is isotropic, and an effective shock speed $u/\cos\theta$ (u : shock speed, θ : shock normal angle) is non-relativistic. The SSDA model cannot consider an anisotropy of the electron pitch-angle distribution. It is well known that the shock with a relativistic effective shock speed has a finite pitch-angle anisotropy. Therefore it is difficult to compare the SSDA model with full particle-in-cell (PIC) simulations which can reproduce only the shock with a relativistic effective shock speed because of the limitation of parameters we can take. To fix this problem, we extend the SSDA model so that it can be applied to such a shock. In this presentation, we discuss the consistency between the extended SSDA model and PIC simulation results.

Keywords: acceleration of particle, Earth, shock waves