

## Constraint of salt compositions on Europa's surface from observations using Subaru telescope and laboratory experiments

\*Shuya Tan<sup>1,2</sup>, Yasuhito Sekine<sup>1</sup>, Masayuki Kuzuhara<sup>3,4</sup>, Taisei Kurosaki<sup>5</sup>, Takashi Kikuchi<sup>5</sup>, Hisayuki Suematsu<sup>5</sup>, Tetsuya Hama<sup>6</sup>, Yoshio Takahashi<sup>2</sup>

1. Earth-Life Science Institute (ELSI), Tokyo Institute of Technology, 2. Department of Earth and Planetary Science, Graduate School of Science, The University of Tokyo, 3. Astrobiology Center, NINS, 4. National Astronomical Observatory of Japan (NAOJ), 5. Extreme Energy-Density Research Institute (EDI), Nagaoka University of Technology, 6. Komaba Institute for Science, The University of Tokyo

The surface of Europa has been recently observed in wavelength of 1.5-2.5  $\mu\text{m}$  with large ground-based telescopes, Keck and VLT (Fischer et al., 2015; Ligier et al. 2016). Those observed reflectance spectra suggest that the existence of chlorides on Europa's geologically active regions, called chaotic terrains. Although those salts would reflect the chemical compositions of the subsurface ocean (e.g., Zolotov et al., 2009), the detailed compositions (i.e., cations and redox state) are not well constrained due to the lack of observations with wide wavelength ranges. Chlorides on the surface could experience dehydration or oxidation to form oxychlorines, such as perchlorates, triggered by irradiations of UV and high-energy particles (e.g., Johnson et al., 2019). However, those are poorly investigated based on laboratory experiments.

Here, we perform high-spectral observations for Europa's surface in the wavelength of 0.95-1.5  $\mu\text{m}$  using the Subaru telescope and IRCS. This provides reflectance spectra of the surface with high-spatial resolution of 200-400 km/pixel, including chaotic terrains. Our results show no remarkable peaks of hydrated salts in the observed spectra. This strongly suggests that surface salts are anhydrous sodium chlorides (e.g., NaCl and/or NaClO<sub>4</sub>).

We also conduct laboratory experiments to simulate irradiations for surface hydrochlorides at low temperatures. We irradiated UV or electron ( $\sim 10$  keV) onto mixtures of H<sub>2</sub>O ice and chlorides (NaCl) at low temperatures ( $\sim 100$  K) using a high-vacuum experimental system. In addition, irradiated higher-energy electron ( $\sim 2.0$  MeV) onto frozen salt solutions with the pulsed power generator "ETIGO-III" (Tokuchi et al., 1998). Our preliminary results show no clear evidence of formation of oxychlorines from NaCl or its mixture with H<sub>2</sub>O ice. Given the dose rate of electron onto Europa's surface (Paranicas et al., 2009), the irradiation doses in the experiments correspond to  $< \sim 100$  years on the surface. These results suggest that little oxychlorines would be formed in chaotic terrains in this short timescale.

Keywords: Europa, Subsurface ocean, Infrared observation, Laboratory experiment, Surface irradiation