

Equilibrium pressure of gas hydrate enclathrated carbon dioxide isotopologues

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Natural CO₂ hydrate was discovered at the Okinawa Trough, and it is expected that CO₂ hydrate exists in the polar ice caps of the Mars. Isotopic fractionation of CO₂ may provide useful information to understand formation processes of gas hydrate. Luzi *et al.* (2011) revealed that CO₂ δ¹³C in hydrate-bound gas is 0.9‰ lower than that of residual gas, suggesting that light CO₂ molecules prefer to be engaged into clathrate cages. We reported at the last JpGU conference (Kimura *et al.*, 2019) that the isotopic difference between GH and residual gases distribute between between 1.2‰ and 1.5‰, agreed fairly well with the previous report. Ozeki *et al.* (2018) reported that the equilibrium pressure of CH₃D is larger than that of CH₄, causes isotopic fractionation in D between GH and residual gases as reported by Hachikubo *et al.* (2007). In this study, we measured the equilibrium pressures of ¹²CO₂ and ¹³CO₂ hydrates and checked the difference between them. CO₂ hydrate samples were synthesized in small pressure cells (volume: 5 mL). Fine ice powder was put in a pressure cell and introduced guest ¹³CO₂ gas. Hydrate crystals were formed by melting the ice powder at the temperature of the quadruple point. We also prepared normal CO₂ (mixture 98.9% ¹²CO₂ and 1.1% ¹³CO₂) hydrate as a reference, using the same preparation method. These pressure cells were placed in a temperature-controlled liquid bath, and measured their equilibrium pressures from 270 K to 278 K. The equilibrium pressure of ¹³CO₂ hydrate was about 0.01 MPa larger than that of normal CO₂ (mainly ¹²CO₂) hydrate, and it agreed with the previous results (Luzi *et al.*, 2011; Ozeki *et al.*, 2018).

References

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