

Thermophysical Properties of Ryugu Estimated from a Box-A Observation

*Yuri Shimaki¹, Hiroki Senshu², Naoya Sakatani¹, Tatsuaki Okada^{1,3}, Tetsuya Fukuhara⁴, Satoshi Tanaka^{1,5}, Makoto Taguchi⁴, Takehiko Arai⁶, Hirohide Demura⁷, Kentaro Suko⁷, Tomohiko Sekiguchi⁸, Toru Kouyama⁹, Jun Takita¹⁰, Sunao Hasegawa¹

1. Institute of Space and Astronautical Science, Japan Aerospace Exploration Agency, 2. Chiba Institute of Technology, 3. University of Tokyo, 4. Rikkyo University, 5. The Graduate University for Advanced Studies, SOKENDAI, 6. Ashikaga University, 7. University of Aizu, 8. Hokkaido University of Education, 9. National Institute of Advanced Industrial Science and Technology (AIST), 10. Hokkaido Kitami Hokuto High School

TIR has acquired many one-asteroid rotation thermal images of Ryugu, during the proximity phase in 2018–2019. High-resolved thermal images were obtained during the Mid-Altitude Observation Campaign on August 1, 2018, with the resolution of ~ 4.5 m/pixel, at the sub-solar latitude of 8°S , the solar phase angle of $\sim 20^\circ$, and the heliocentric distance of 1.06 AU. The diurnal temperature profiles of Ryugu showed flat patterns. A thermophysical model using a shape model of rough surface well reproduced the diurnal temperature profiles. By comparing the observation and calculation results, the global thermal inertia of Ryugu was estimated to be $225 \pm 45 \text{ J m}^{-2} \text{ s}^{-0.5} \text{ K}^{-1}$, and the global surface roughness was determined as 0.41 ± 0.08 (Shimaki et al., submitted to *Icarus*). Because of the sub-solar latitude, we cannot determine the thermophysical properties in a part of the northern hemisphere. Here we report the thermophysical properties of Ryugu determined by a Box-A observation on November 14, 2018, with the resolution of ~ 20 m/pixel, at the sub-solar latitude of 2°S , the solar phase angle of 4.8° , and the heliocentric distance of 1.35 AU. The diurnal temperature profiles showed a peak around the late afternoon. We confirmed that the thermophysical model well reproduces the observed temperature profiles. The global distribution of the thermal inertia is similar to that determined by the previous study, however, we see an offset of the values and the global thermal inertia was estimated to be $\sim 300 \text{ J m}^{-2} \text{ s}^{-0.5} \text{ K}^{-1}$. The global surface roughness was estimated to be ~ 0.3 , with the small values around the equatorial ridge.

Acknowledgments: We would like to thank the Hayabusa2 team members for their technical and operational support and helpful scientific discussions. This work is partly supported by the JSPS Grants-in-Aid for Scientific Research (KAKENHI JP17H06459 and JP19H01951) and the JSPS Core-to-Core program “International Network of Planetary Sciences” .

Keywords: Hayabusa2, Thermal Infrared Imager, Asteroid Ryugu