

## Current production rate of dark splotches on the surface of Mars

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High-resolution images of Mars surface have revealed that some small-scale features are currently changing/occurring in numerous locations. Those include moving dunes, recurring slope lineae [McEwen *et al.*, 2011]), slope streaks [Malin and Edgett, 2001], gullies and their associated rockslides [Malin *et al.*, 2006]), and 10' s to 100' s meter-scale dark splotches, which are commonly interpreted as results of small impacts [Malin *et al.*, 2006]. The production rate of such splotches (or dark patches) were estimated through image analyses [Malin *et al.*, 2006; Daubar *et al.*, 2013] and from the current impact flux [JeongAhn and Malhotra, 2015; Williams *et al.*, 2014].

The approach of Daubar *et al.* [2013], carefully comparing high-resolution images of the same location at different times, is the most straight-forward but requiring significant effort of humans because automatic machine comparisons of different images obtained at different illumination and calibration conditions are technically difficult. To overcome this difficulty, we perform thorough comparisons of images by using >20,000 people in a museum located at the center of Tokyo. More than 2,450 pairs of CTX images (~3.2 years apart on average), equivalent to the surface areas of  $\sim 6.2 \times 10^6 \text{ km}^2$  (ATF =  $\sim 2.0 \times 10^7 \text{ km}^2 \text{ yr}$ ), were manually analyzed through about 15 months of the experiment. The results are also examined by 20 volunteer members of the museum, and then reanalyzed by 4 researchers. The entire team is organized by three professional scientists of Mars geology background. So far, we identify >230 newly formed splotches including five reported in [Daubar *et al.*, 2013]). We obtain the production rate of  $\sim 1.1 \times 10^{-5} / \text{km}^2 / \text{yr}$ , which is more than one order of magnitude larger than the values reported in previous image-based studies [Malin *et al.*, 2006; Daubar *et al.*, 2013] and theoretical models [Hartmann, 2005; JeongAhn and Malhotra, 2015]. In this talk, we will also report the arial difference of the production rate.

### References

- Daubar, I. J., A. S. McEwen, S. Byrne, M. R. Kennedy, and B. Ivanov (2013), The current martian cratering rate, *Icarus*, 225(1), 506-516.
- Hartmann, W. K. (2005), Martian cratering 8: Isochron refinement and the chronology of Mars, *Icarus*, 174 (2), 294-320.
- JeongAhn, Y., and R. Malhotra (2015), The current impact flux on Mars and its seasonal variation, *Icarus*, 262, 140-153.
- Malin, M. C., and K. S. Edgett (2001), Mars Global Surveyor Mars Orbiter Camera: Interplanetary cruise through primary mission, *Journal of Geophysical Research: Planets*, 106(E10), 23429-23570.
- Malin, M. C., K. S. Edgett, L. V. Posiolova, S. M. McColley, and E. Z. N. Dobra (2006), Present-Day Impact Cratering Rate and Contemporary Gully Activity on Mars, *Science*, 314(5805), 1573.
- Malin, M. C., *et al.* (2007), Context Camera Investigation on board the Mars Reconnaissance Orbiter, *Journal of Geophysical Research: Planets*, 112(E5), n/a-n/a.
- McEwen, A. S., L. Ojha, C. M. Dundas, S. S. Mattson, S. Byrne, J. J. Wray, S. C. Cull, S. L. Murchie, N. Thomas, and V. C. Gulick (2011), Seasonal flows on warm Martian slopes, *Science*, 333(6043), 740-743.
- McEwen, A. S., *et al.* (2007), Mars Reconnaissance Orbiter's High Resolution Imaging Science Experiment (HiRISE), *Journal of Geophysical Research: Planets*, 112(E5), n/a-n/a.
- Williams, J.-P., A. V. Pathare, and O. Aharonson (2014), The production of small primary craters on Mars and the Moon, *Icarus*, 235, 23-36.

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