Variations of lo's volcanic activity seen in Jupiter's extended sodium nebula

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Io, which is one of the Galilean moons of Jupiter, is the most volcanically active body in the Solar System. Volcanic atmosphere is ionized and picked-up by Jupiter's co-rotating magnetic fields. This plasma distributes in Jupiter's inner magnetosphere and forms a structure called Io plasma torus. Major constituents in the torus are sulfur and oxygen ions, and most of these ions have emissions lines at UV wavelengths. Although this is a minor constituent, NaCl+ should be included in the torus since Cl+ ions was detected from the ground, and neutral sodium atoms show the most distinct emission at sodium D-line wavelengths in the torus. Not only in the torus, sodium emission can be observed also in a vast region whose extent is 1,000 Jupiter's radii around Jupiter. This structure is called Jupiter's sodium nebula, or Mendillo-sphere. This means these sodium atoms have enough velocity to escape from Jupiter's and lo's gravitational-spheres.

These sodium atoms seem to be originated from sodium chloride in lo's volcanic gas. This gas becomes lo's ionospheric plasma. Pick-up of these NaCl+ ions from lo's ionosphere and their subsequent destruction in the plasma torus produces fast from of neutral sodium atoms, then Jupiter's sodium nebula is formed. This sodium nebula can be observed from the ground using small telescopes.

We have been making observations of Jupiter's sodium nebula atop Heleakala in Maui island, Hawaii, and found the nebula shows variations that seem to correspond to those in lo's volcanic activity. Since 2013, we have been making the observations in conjunctions with the Hisaki and Juno spacecraft. The most distinct event during this campaign was seen in 2015. Other than this, several minor enhancements were observed.

In this presentation, we will show latest sodium data that is representative of lo's volcanism. Also, comparisons of the sodium data with Hisaki's torus data and lo's infrared observations will be shown.

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